

CGSIC, Miami, 9/16/2019

Development of a Satellite-Based Cold Atom Clock

Liang Liu

Key Laboratory of Quantum Optics
Shanghai Institute of Optics and Fine Mechanics
Chinese Academy of Sciences

OUTLINE

1. Introduction

2. Laser cooling in diffuse light

4. Compact cold atom clock

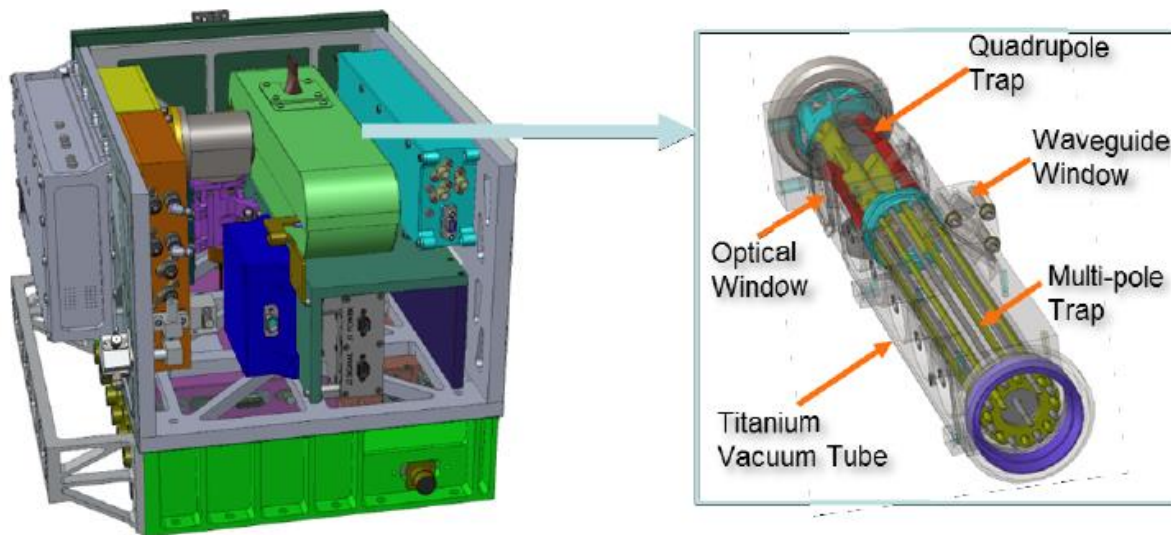
5. Conclusions



OUTLINE

1. Introduction





DSAC DU with the mercury ion trap configuration
29x26x23 cm, 16 kg, 50W, 1E-15/d



What we need for satellite-based atomic clock:

- ❑ Small size: $<300 \times 300 \times 300$ mm
- ❑ Light weight: <20 kg
- ❑ Less power: <50 W
- ❑ Good performance:
 - 1s: $<2E-13$
 - 1day: $<1E-15$
 - Drift: $<1E-15$
 - Uncertainty: $<1E-15$

Besides Mercury ion clock, an atomic clock with **laser cooled atoms** can reach the performance for next generation GNSS and deep space exploration.

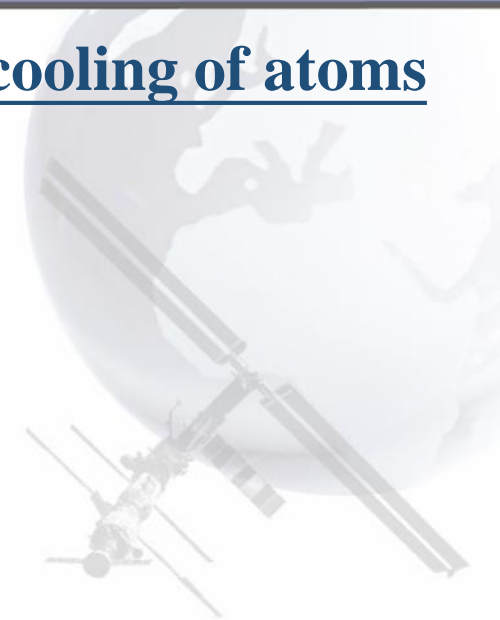
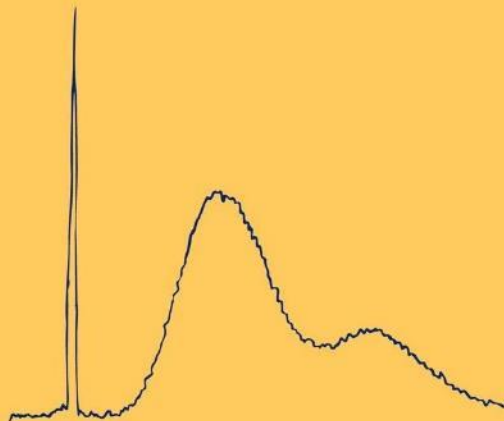


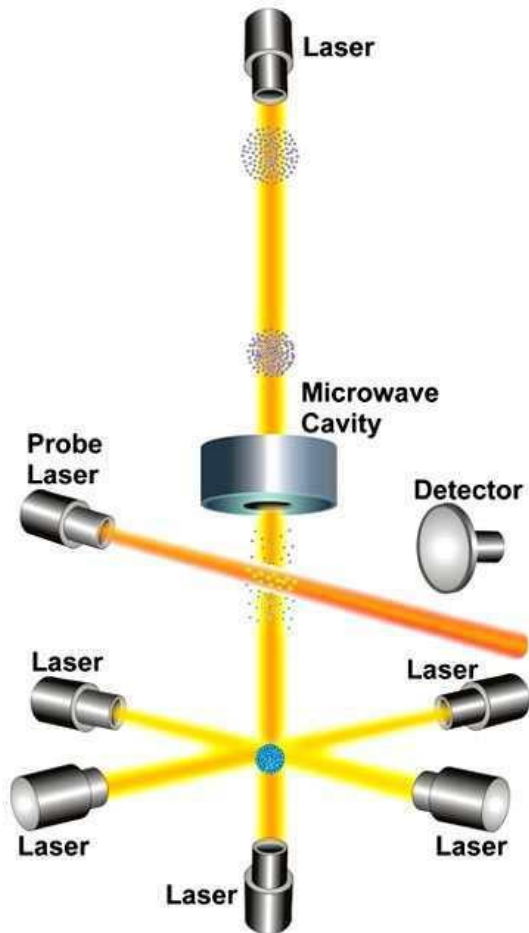


Laser cooling of atoms

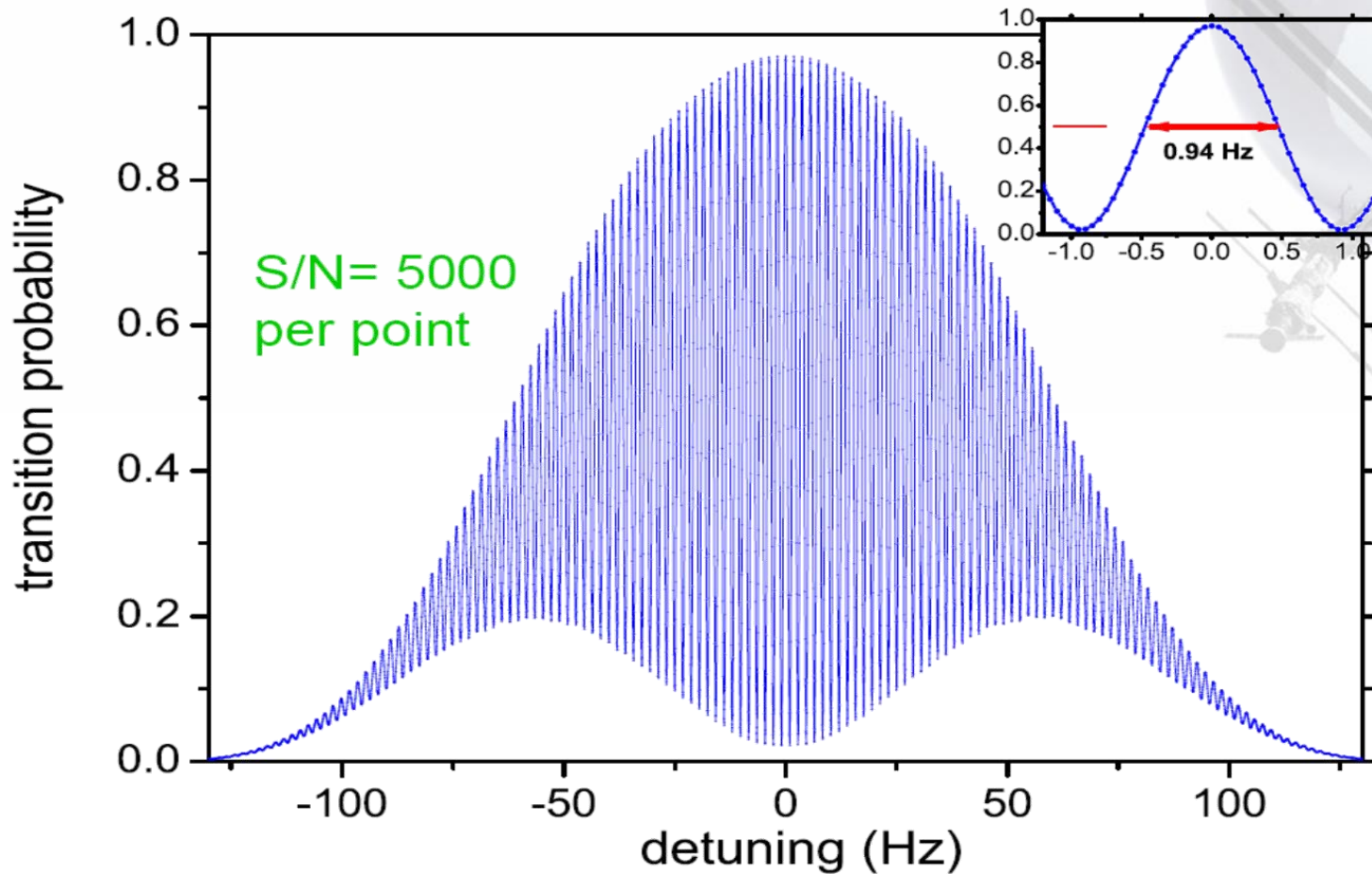
Harold J. Metcalf
Peter van der Straten

Laser Cooling and Trapping





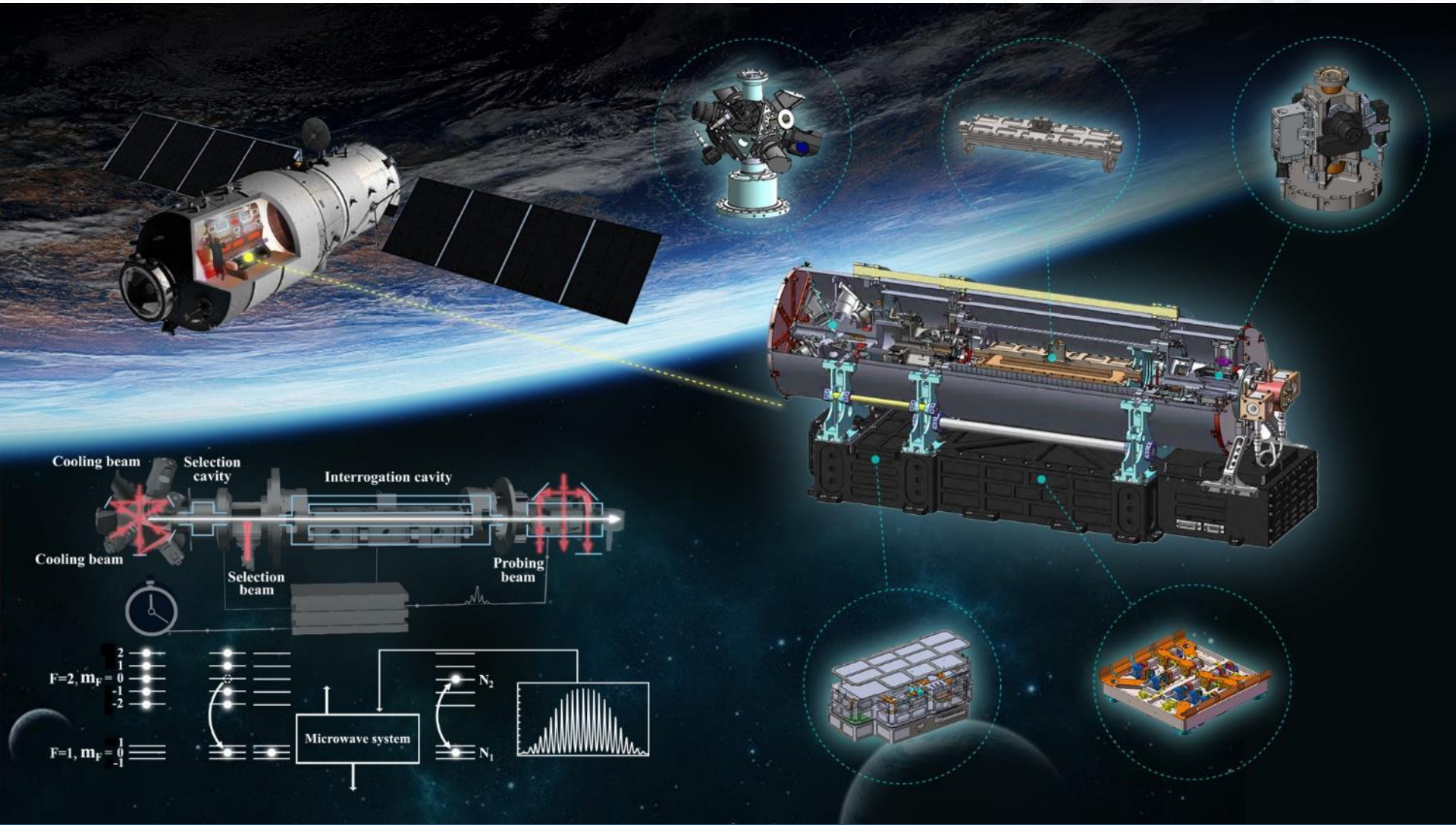
- ❑ Atoms are cooled at the bottom
- ❑ Cold atoms are launched upwards by lasers
- ❑ Cold atoms interact with microwave in the cavity
- ❑ Cold atoms drop after they reach the top
- ❑ Cold atoms interact with microwave again
- ❑ Cold atoms are detected by laser
- ❑ The signal feedback to microwave source



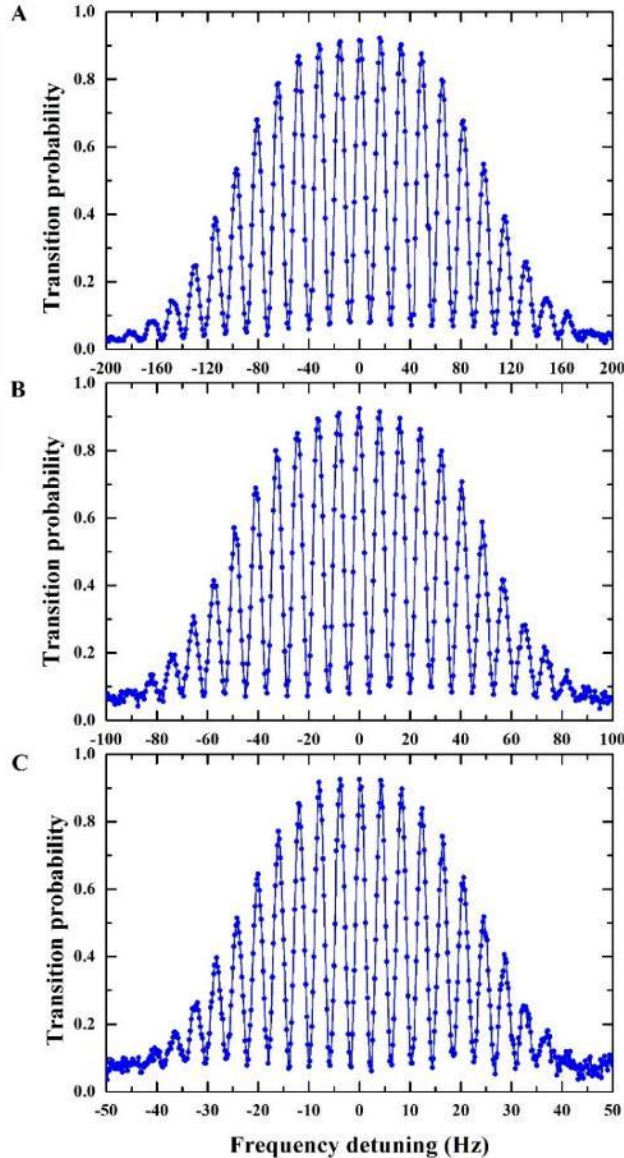
Typical Ramsey fringes, Paris

Systematic fractional frequency shifts for FO1, FO2 and FOM ^{133}Cs fountains

	FO1 ($\times 10^{16}$)	FO2 ($\times 10^{16}$)	FOM ($\times 10^{16}$)
Quadratic Zeeman effect	1199.7 ± 4.5	1927.3 ± 0.3	351.9 ± 2.4
Blackbody radiation	-162.8 ± 2.5	-168.2 ± 2.5	-191.0 ± 2.5
Collisions and cavity pulling	-197.9 ± 2.4	-357.5 ± 2.0	-34.0 ± 5.8
Microwave spectral purity & leakage	0.0 ± 3.3	0.0 ± 4.3	0.0 ± 2.4
First order Doppler effect	< 3	< 3	< 2
Ramsey & Rabi pulling	< 1	< 1	< 1
Microwave recoil	< 1.4	< 1.4	< 1.4
Second order Doppler effect	< 0.08	< 0.08	< 0.08
Background collisions	< 1	< 1	< 1
Total uncertainty	± 7.5	± 6.5	± 7.7



Cold Atom Clock Experiment in Space, CACES, with Tiangong-2



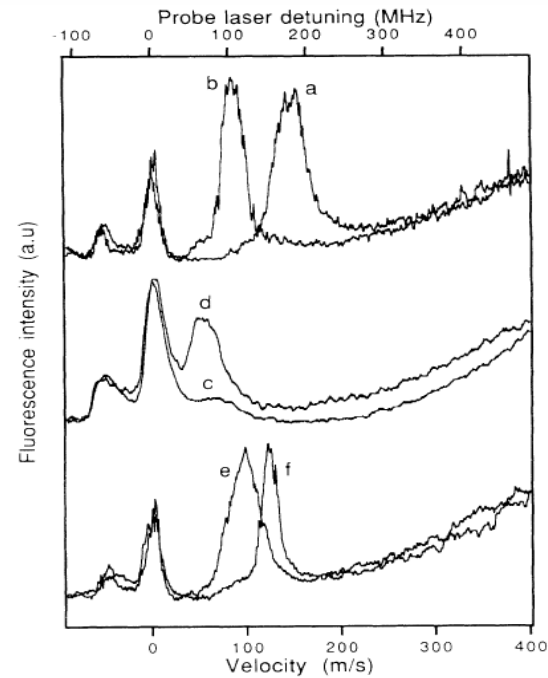
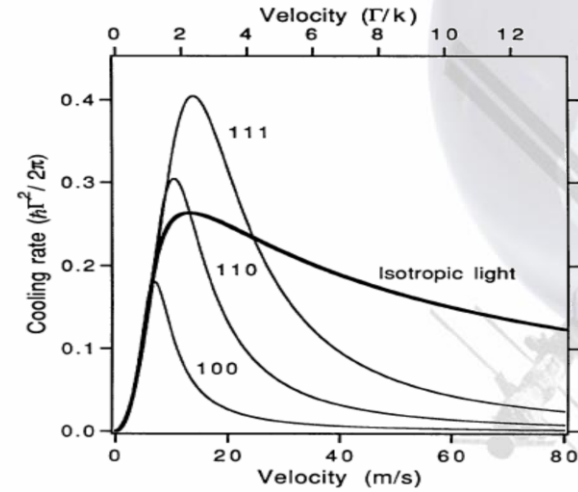
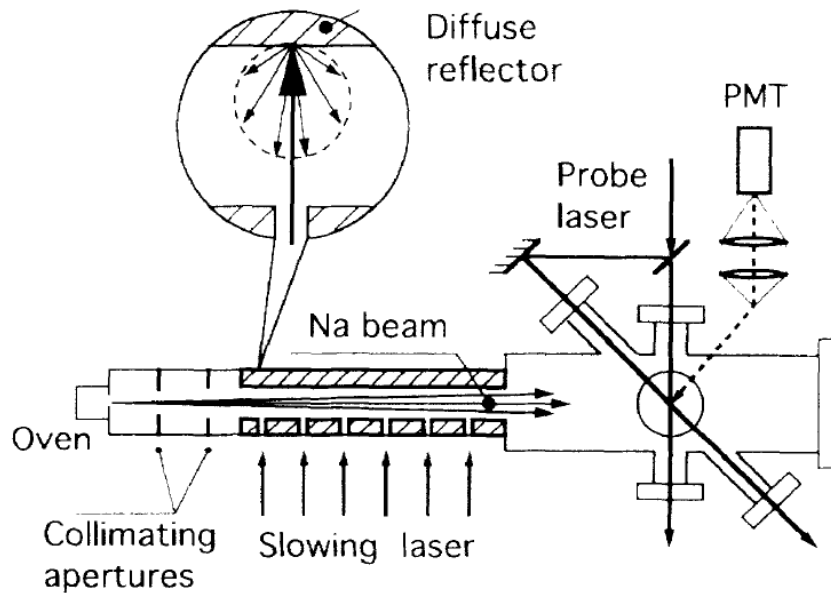
Typical Ramsey fringe. The launching velocity is (A) 3.15 m/s, (B) 1.69 m/s, (C) 0.78 m/s, corresponding to the width (A) 7.27 Hz, (B) 3.89 Hz, and (C) 1.80 Hz.

OUTLINE

2. Laser cooling in diffuse light



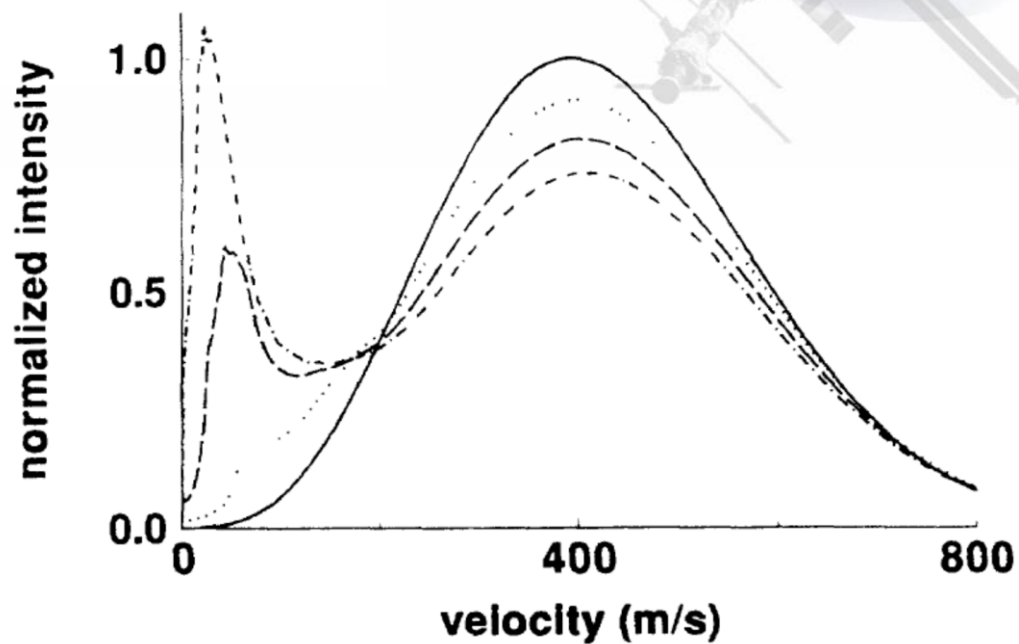
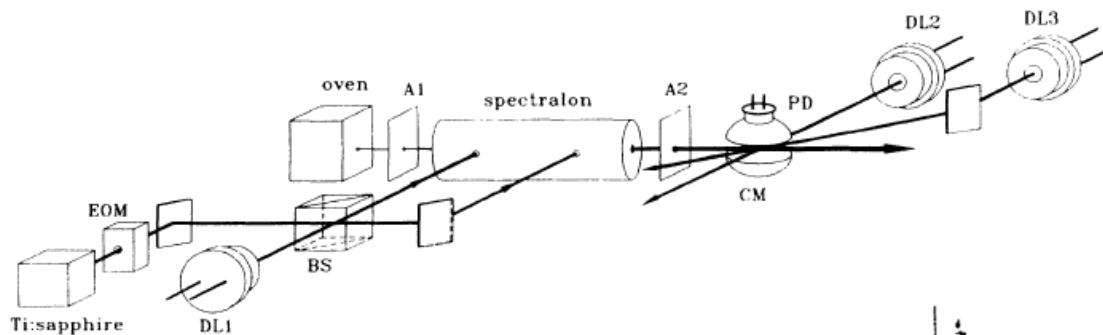
$$\omega_a = \omega_1 + kv \cos\theta$$



W. Ketterle, et. al., Phys. Rev. Lett. 69, 2483 (1992)

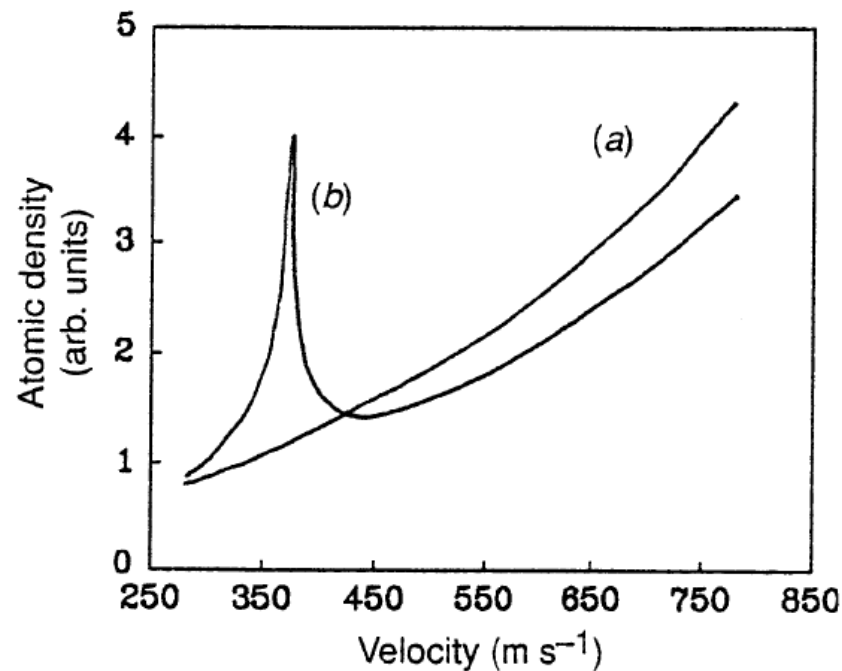
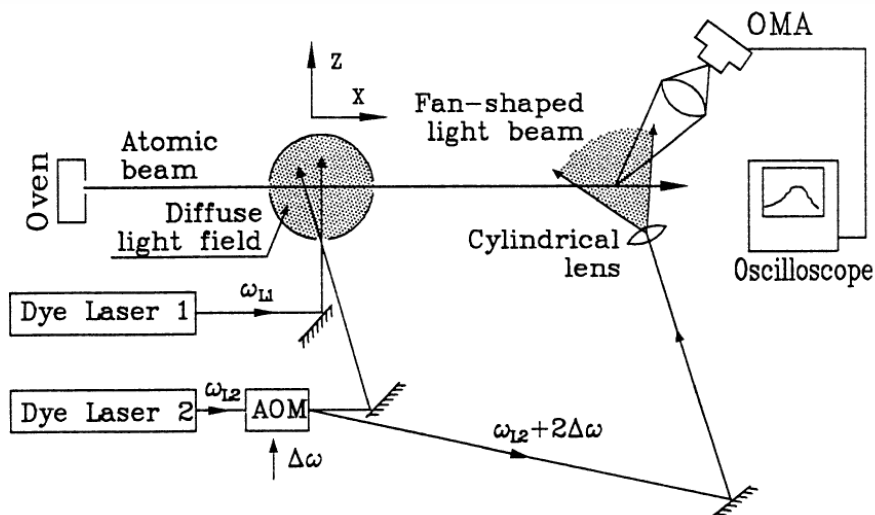


Diffuse deceleration of an atomic beam



H. Batelaan, et.al., Phys. Rev. A 49, 2780 (1994)





H.X. Chen, L. Liu and Y.Z. Wang, Acta Optica Sinica 14, 125 (1994)

Y.Z. Wang and L. Liu, Australian J. Phys. 48, 267 (1995)

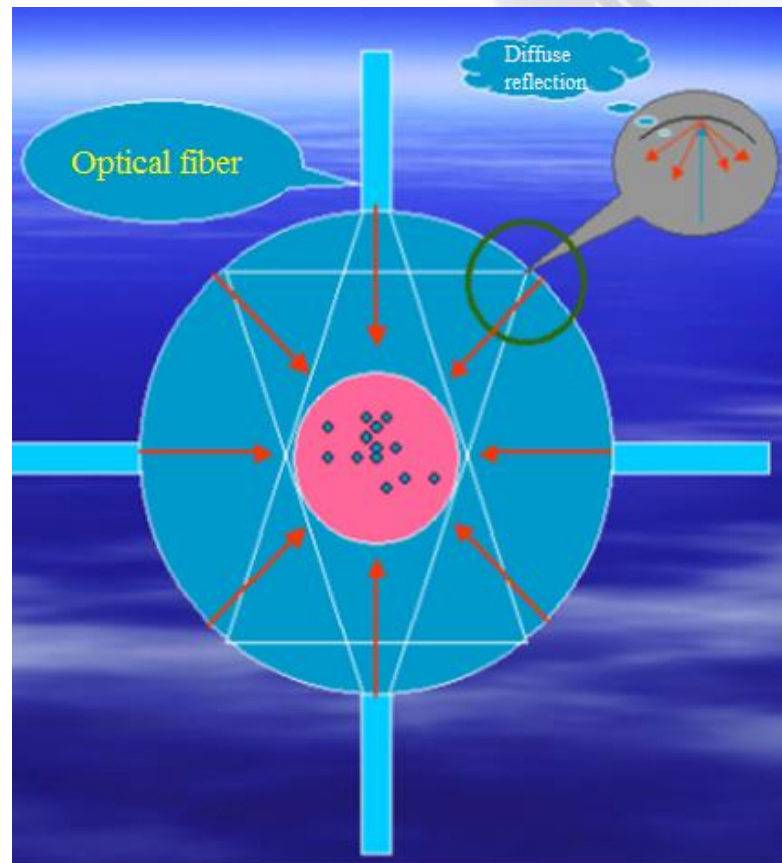


Laser cooling in an integrating sphere

A new scheme of laser cooling is introduced in order to reduce the device size:
diffuse laser cooling

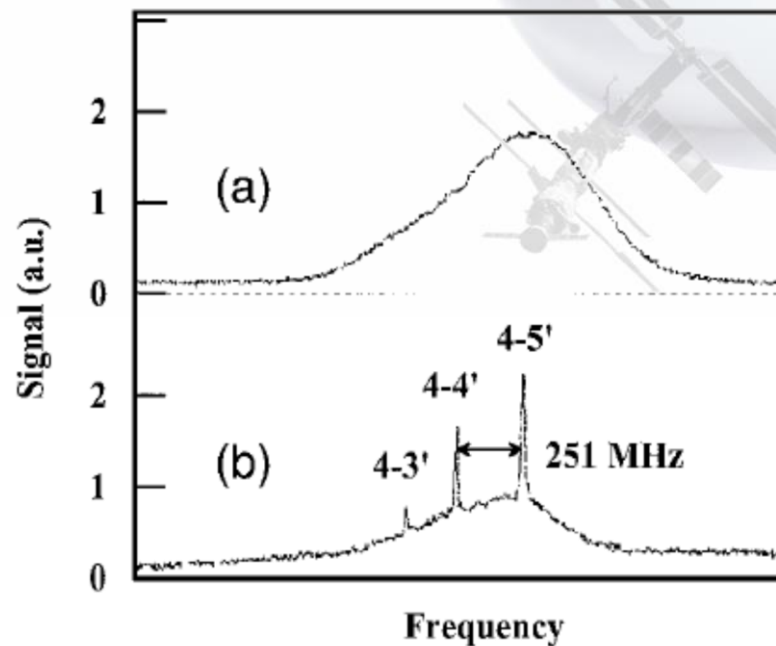
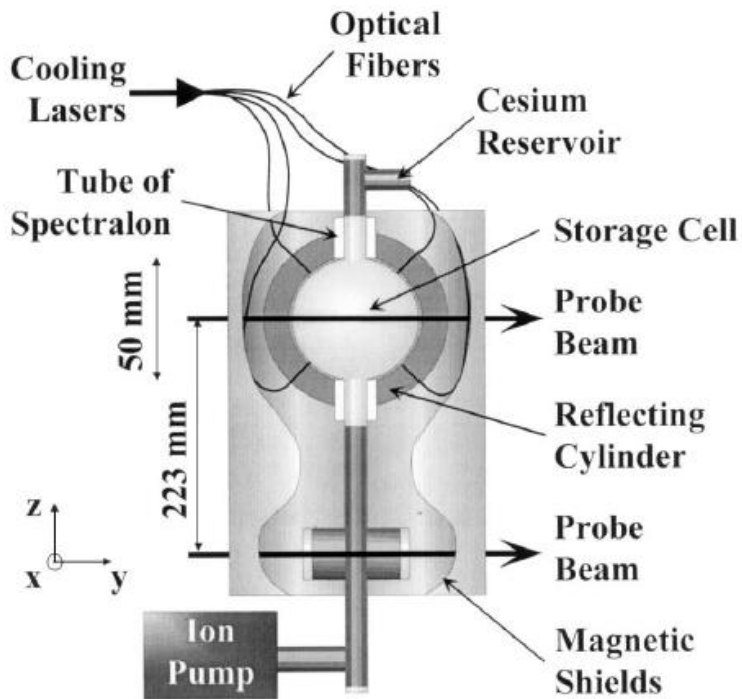


Typical integrating sphere

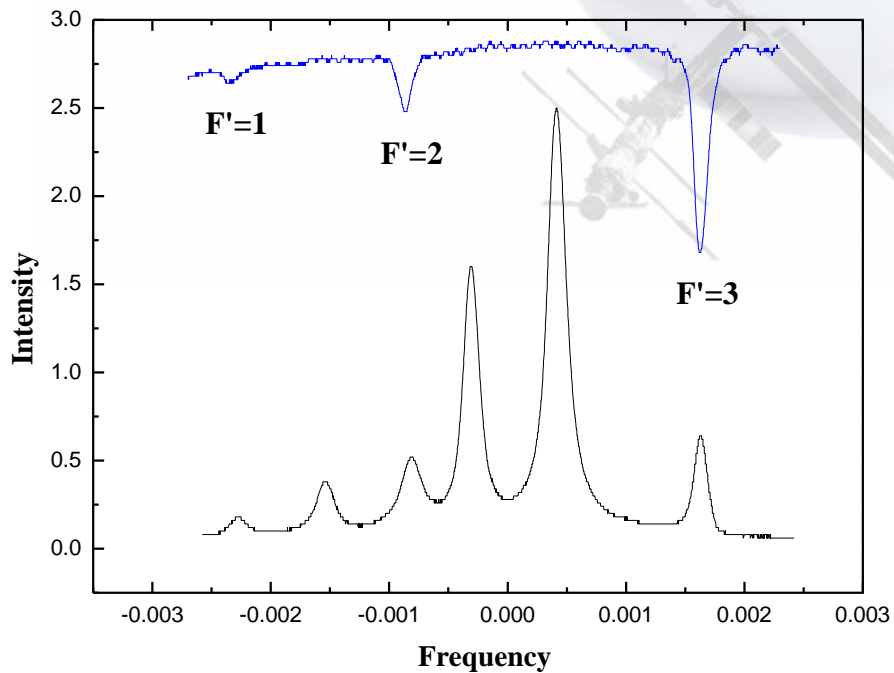
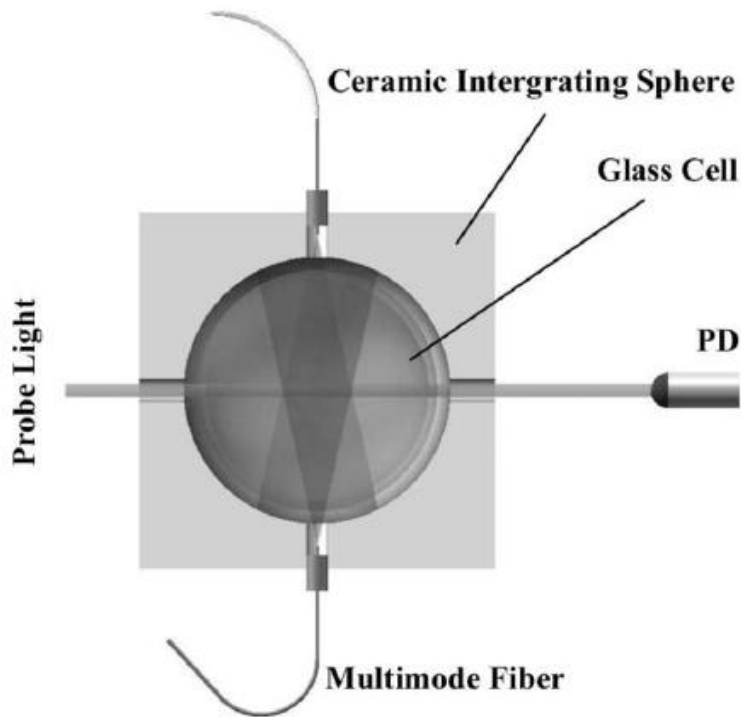


Principle of laser cooling in an integrating sphere





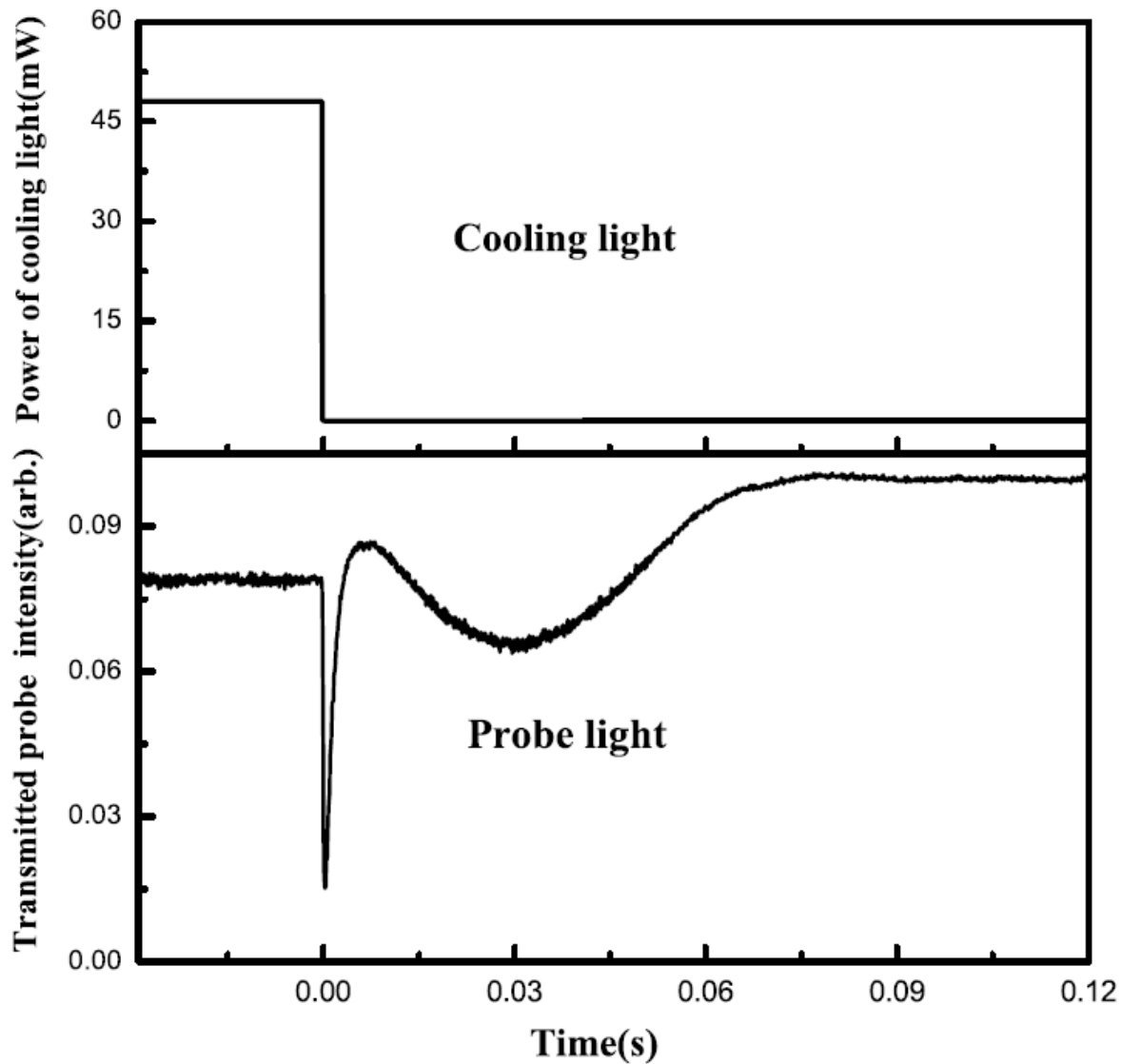
E. Guillot, *et al.*, Opt. Lett. 26, 1639 (2001), Cs



H.D. Cheng, *et al.*, Phys. Rev. A 79, 023407 (2009), Rb

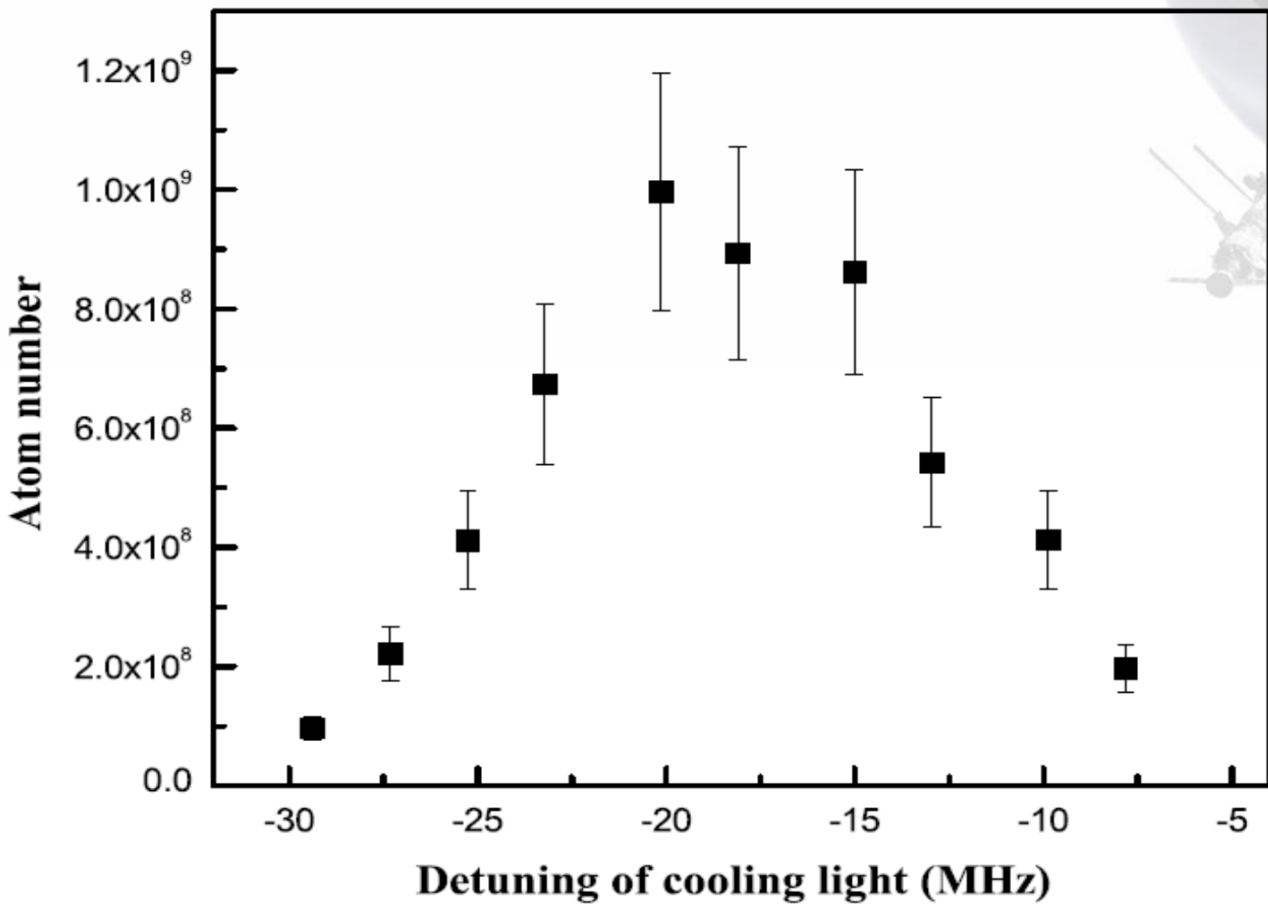


Laser cooling of atoms in diffuse light



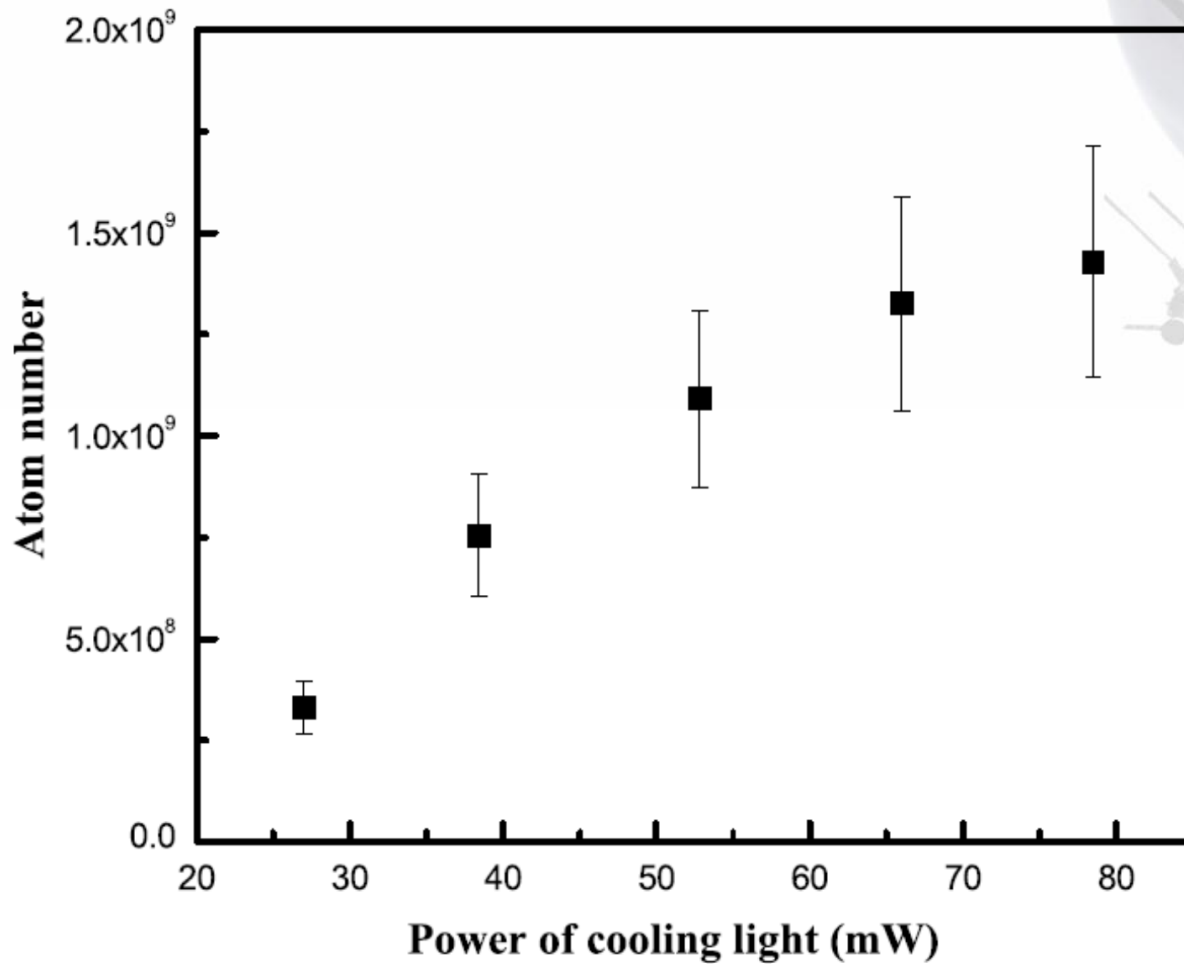


Laser cooling of atoms in diffuse light

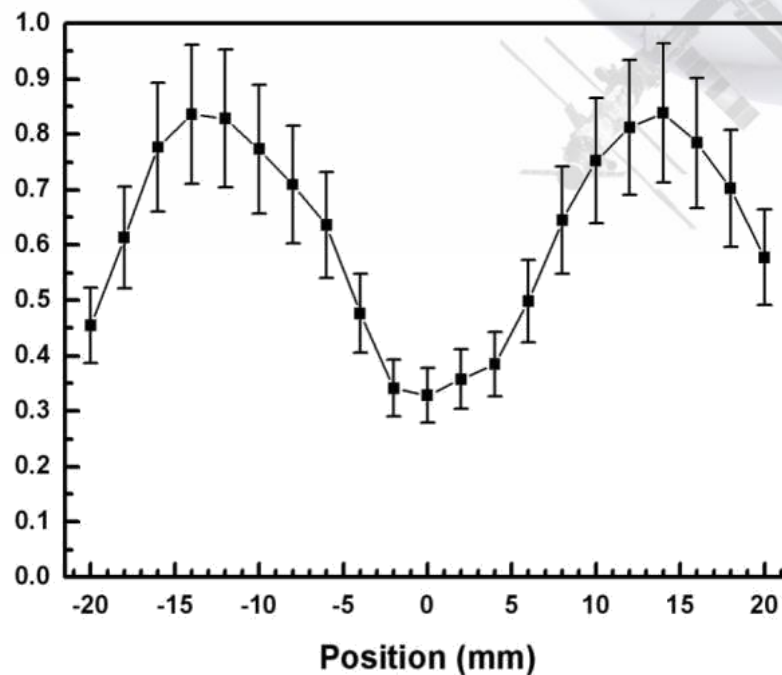
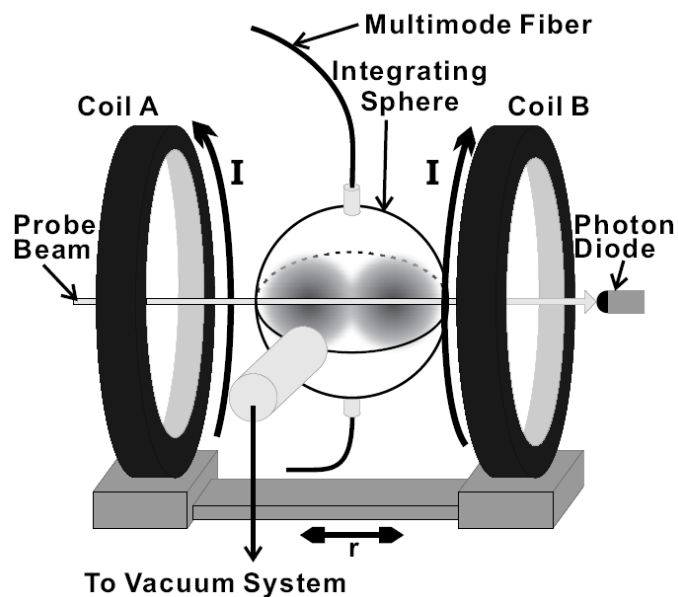


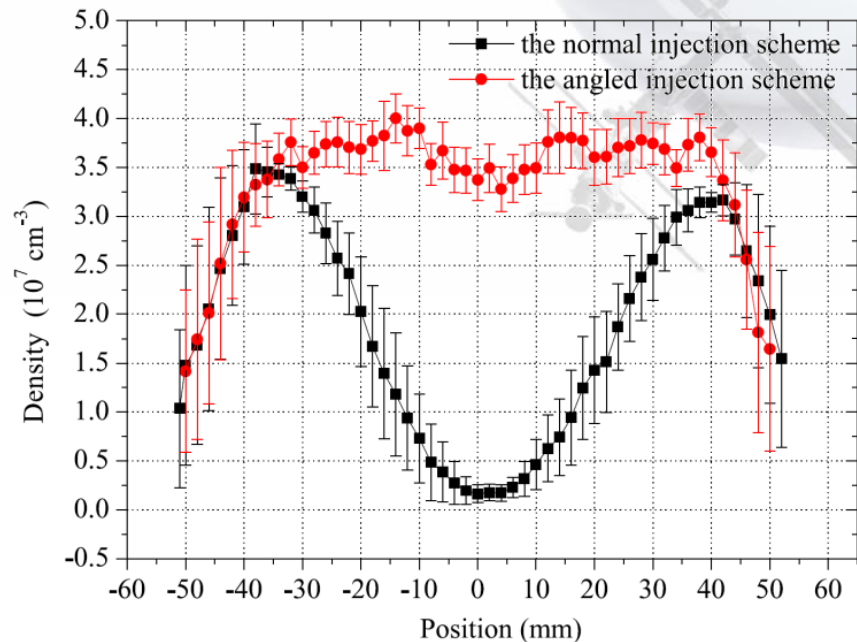
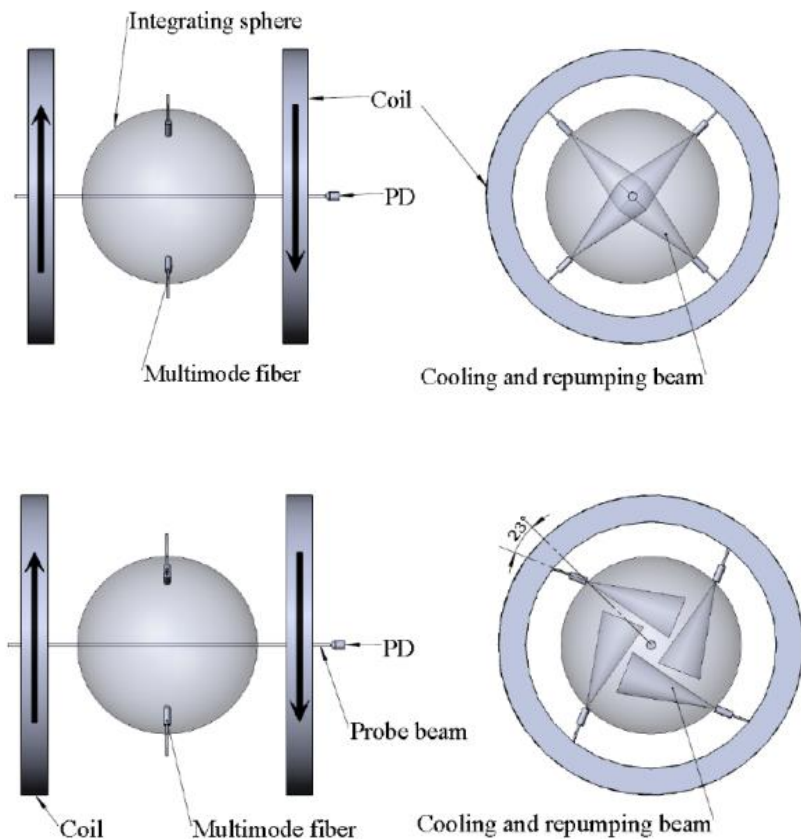


Laser cooling of atoms in diffuse light



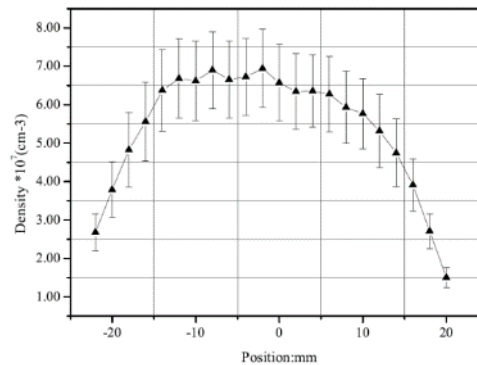
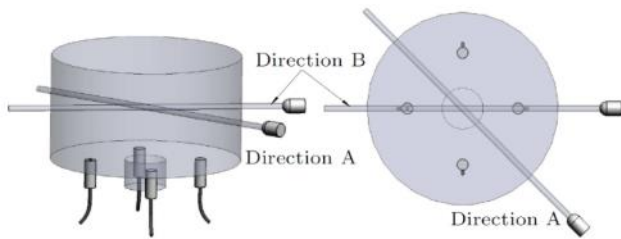
Controlling the distribution of atomic density



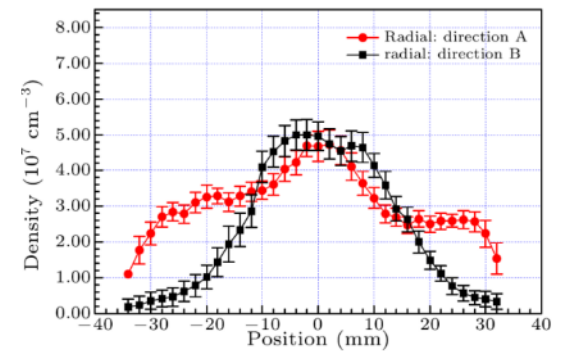


Controlling distribution of cold atom density in a **spherical** cavity

Controlling the distribution of atomic density



Axial



Radial

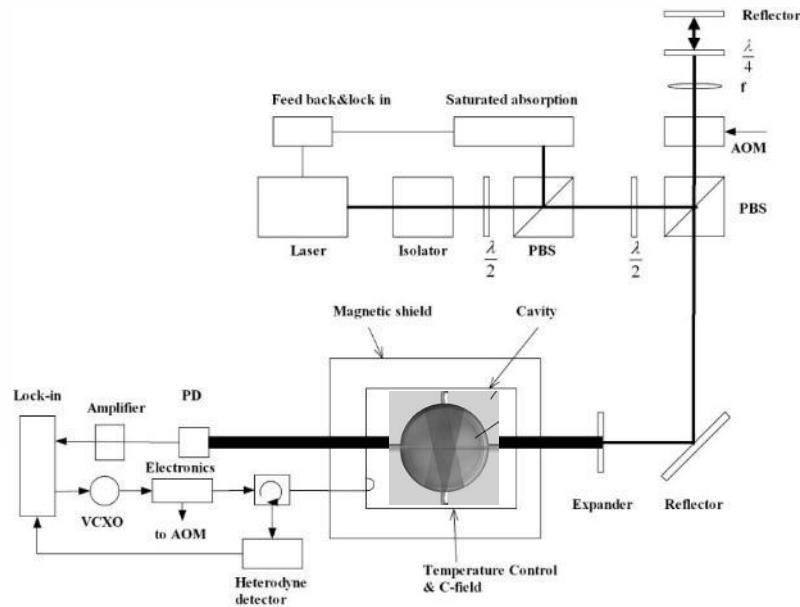
Controlling distribution of cold atom density in a **cylindrical** cavity

OUTLINE

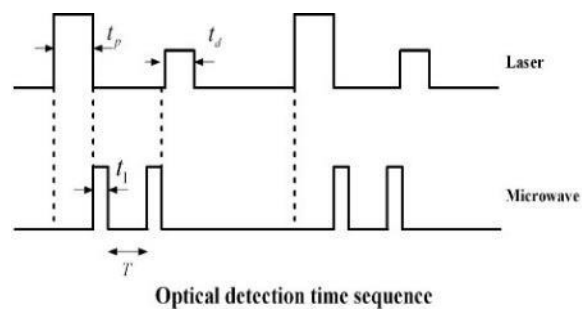
3. Compact cold atom clock



Atomic clock with diffuse laser cooled atoms



A Pulsed Optically Pumped (POP) scheme: optical pumping, interrogation and detection are separate to avoid light shift.



A Pulsed Optically Pumped (POP) scheme with diffuse laser cooled atoms in an integrating sphere (ISCAC)

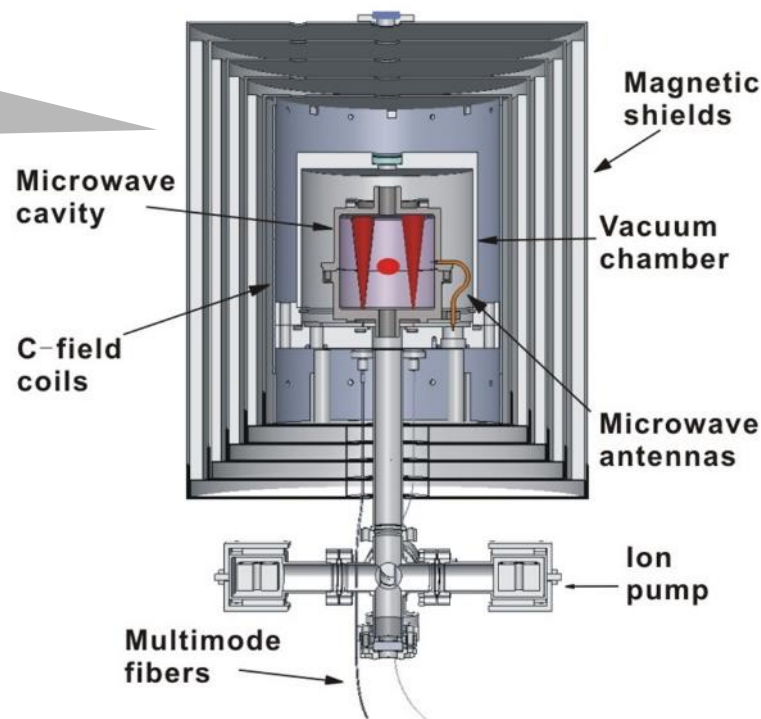
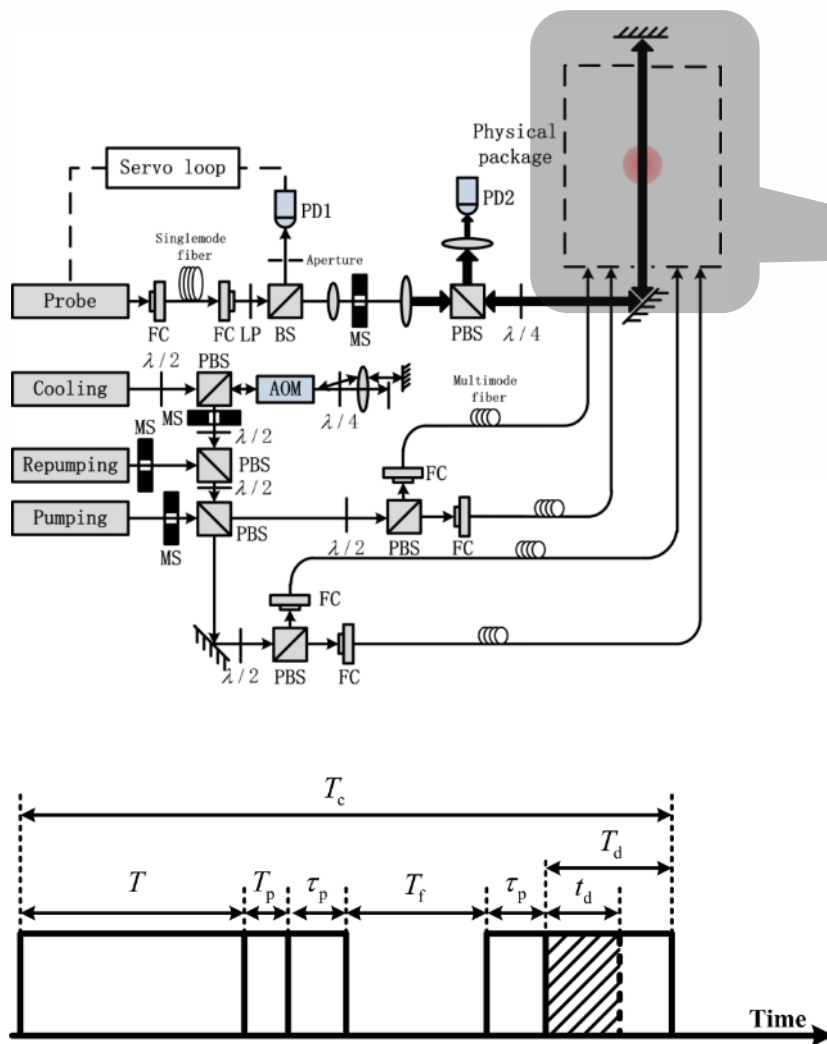


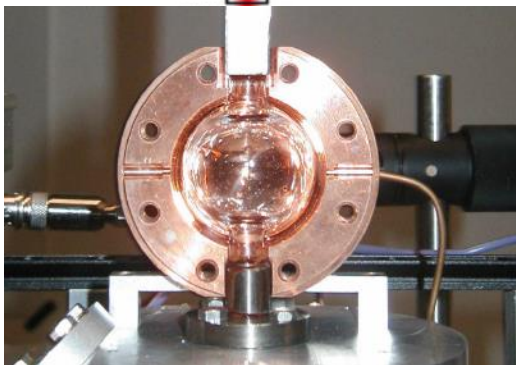
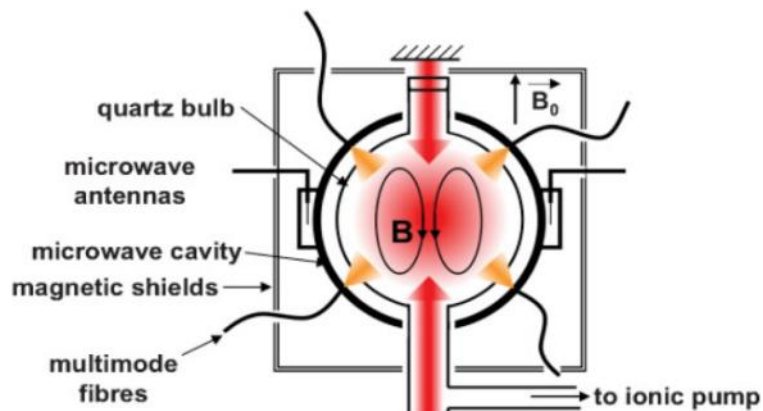
	Cs	Rb
Atom	^{133}Cs	^{87}Rb
Microwave frequency	9.192632 GHz	6.834682 GHz
D2 line	F=2	F=1
Melting temperature	28.84 °C	38.89 °C
Cooling laser wavelength	852.356 nm	780.241 nm



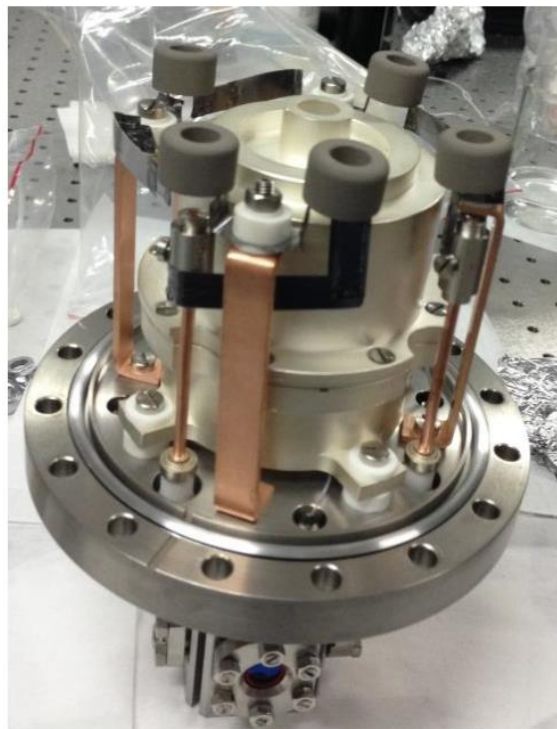


Atomic clock with diffuse laser cooled atoms

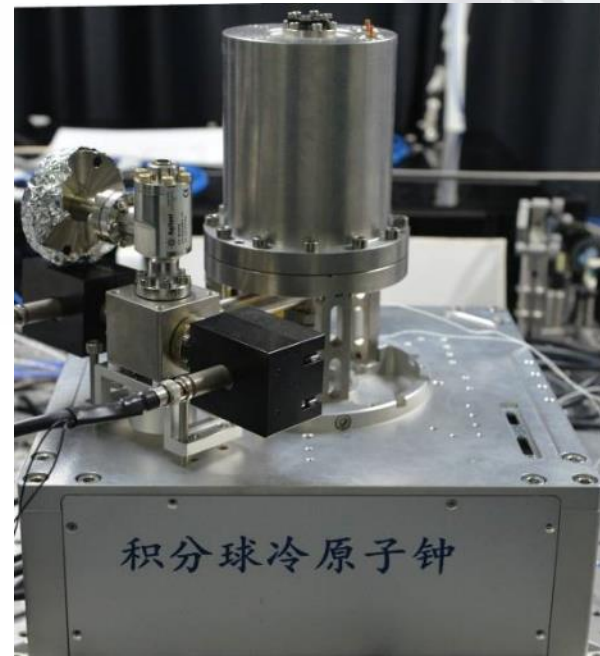
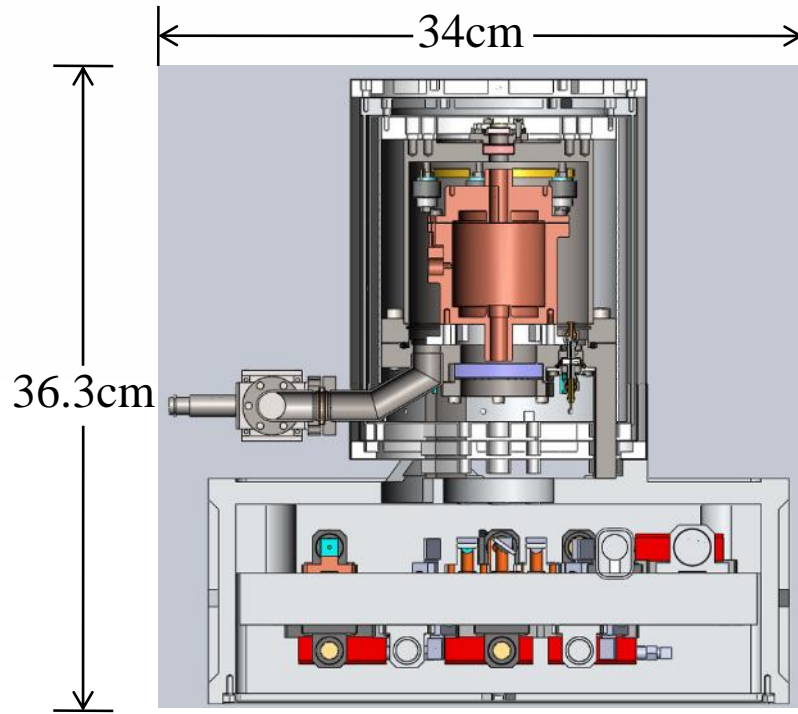




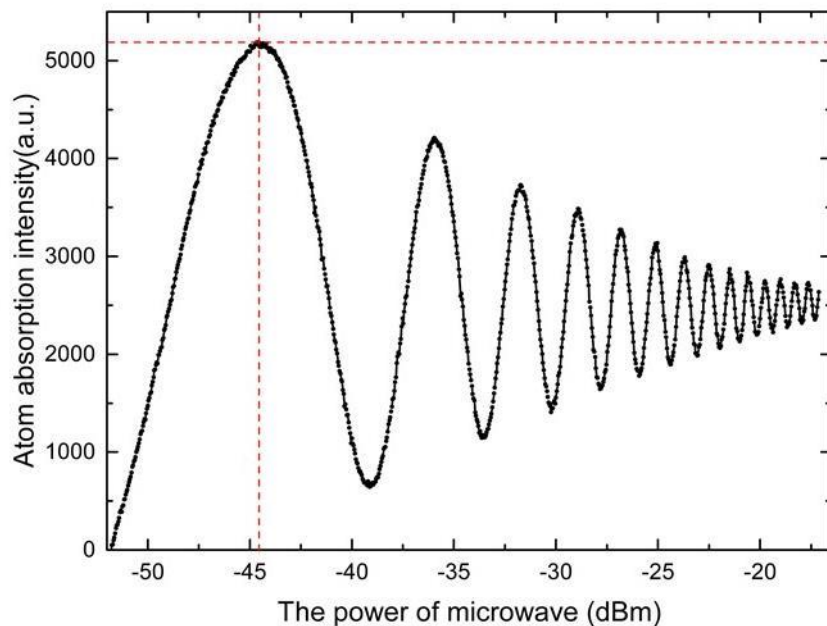
HORACE, Paris



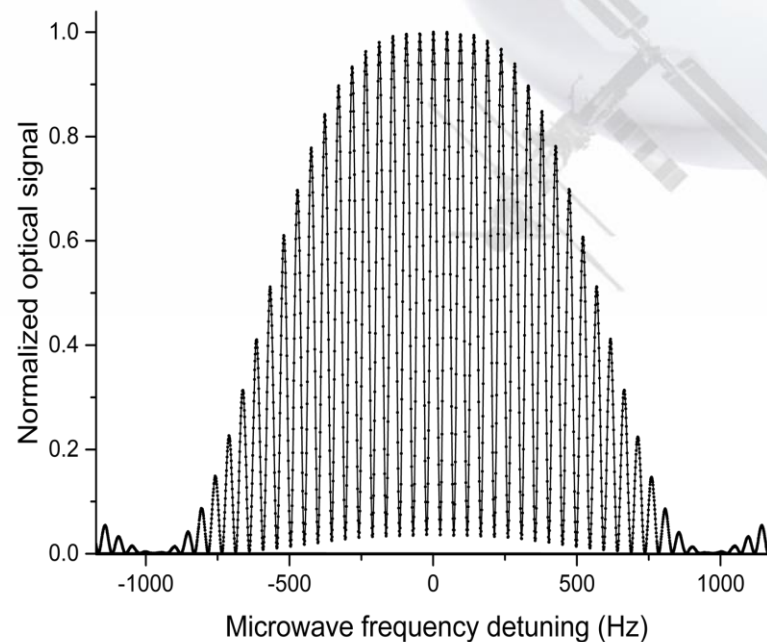
ISCAC, Shanghai



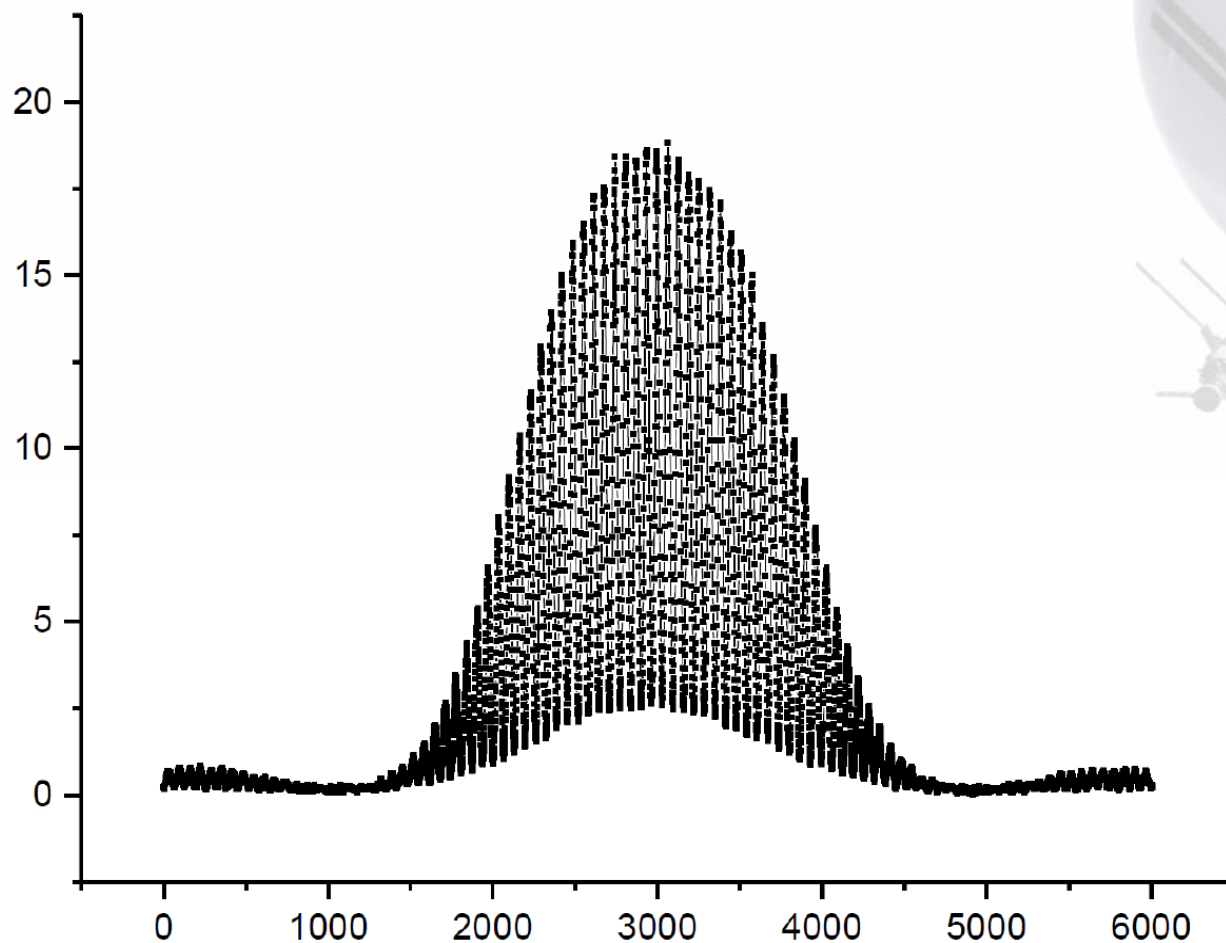
Engineering model



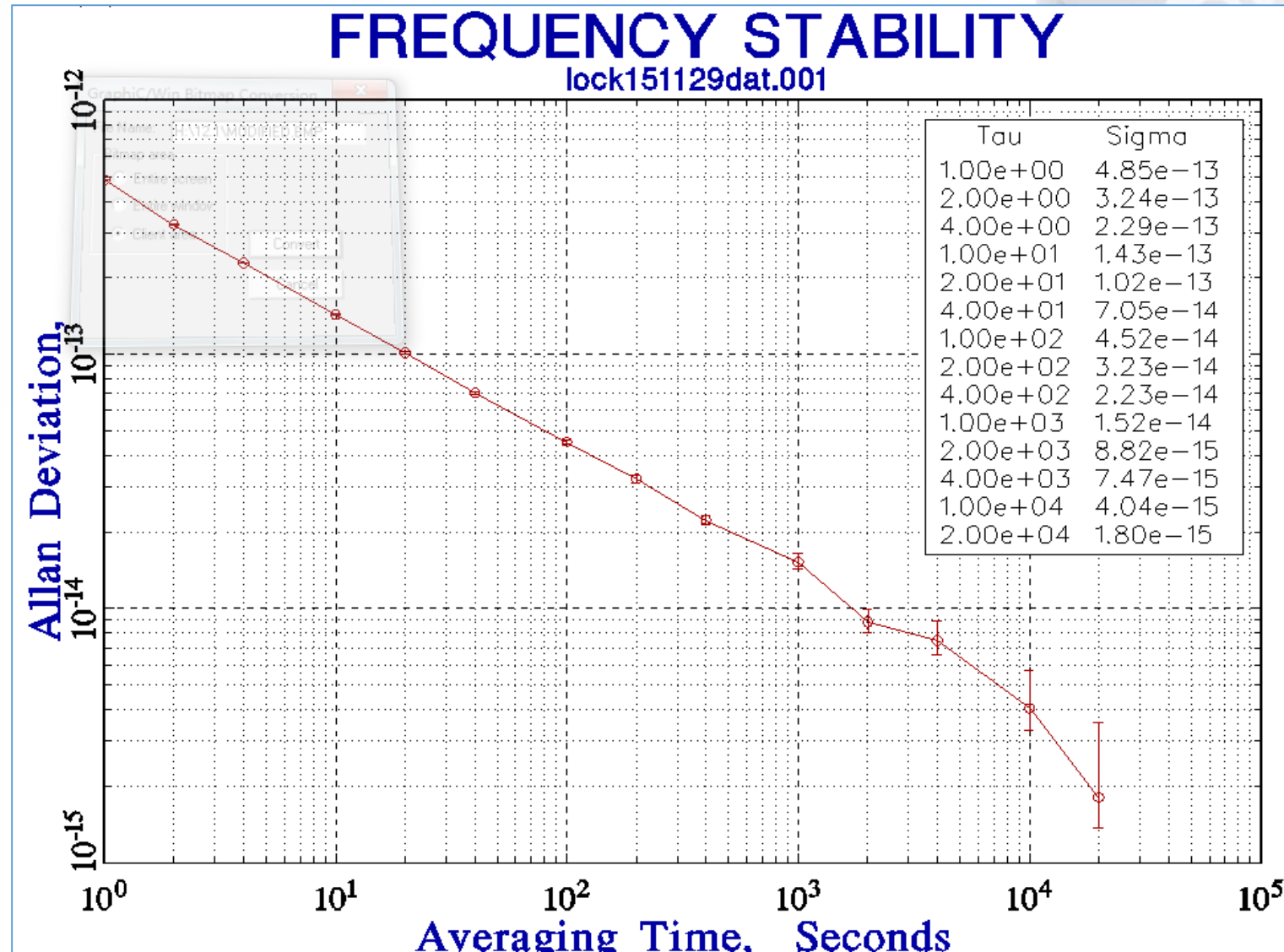
Rabi oscillation

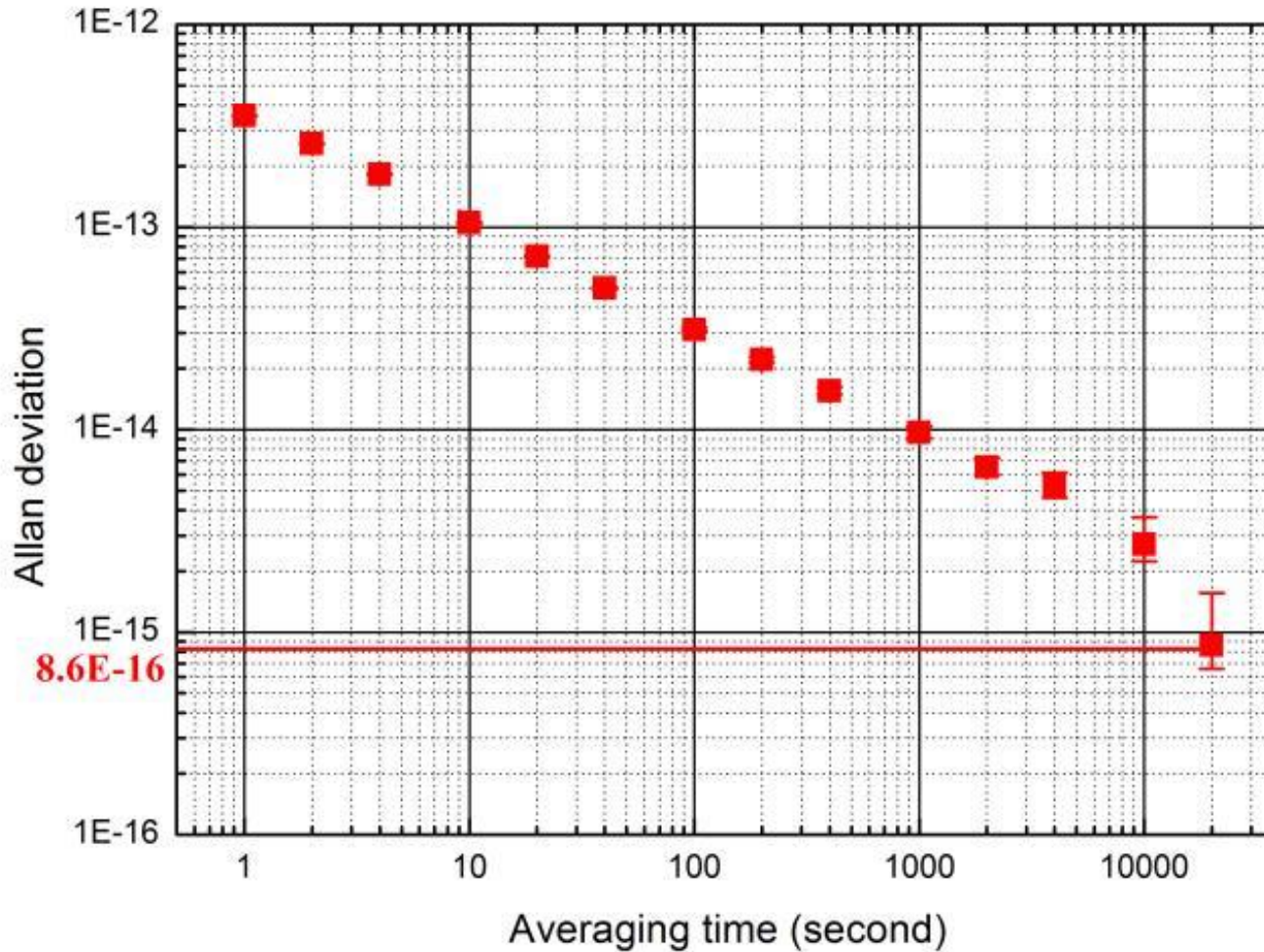


Ramsey fringe with linewidth 20 Hz



**Ramsey fringe with
linewidth 13.5 Hz**



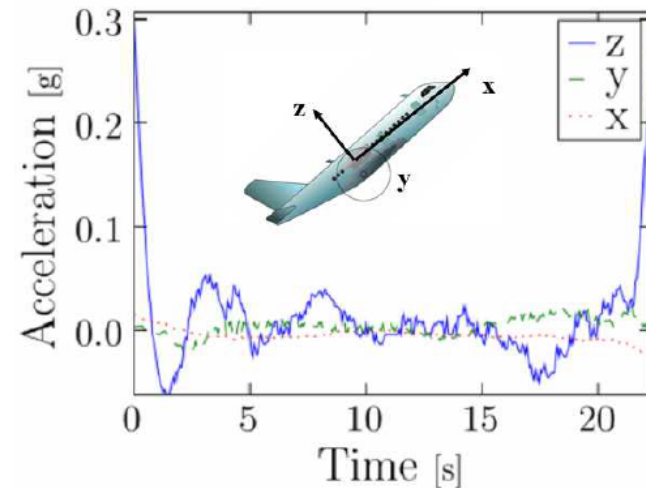
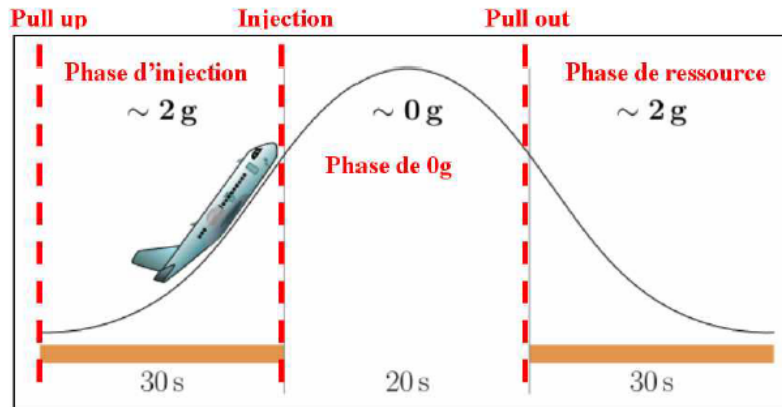


OUTLINE

4. Conclusions



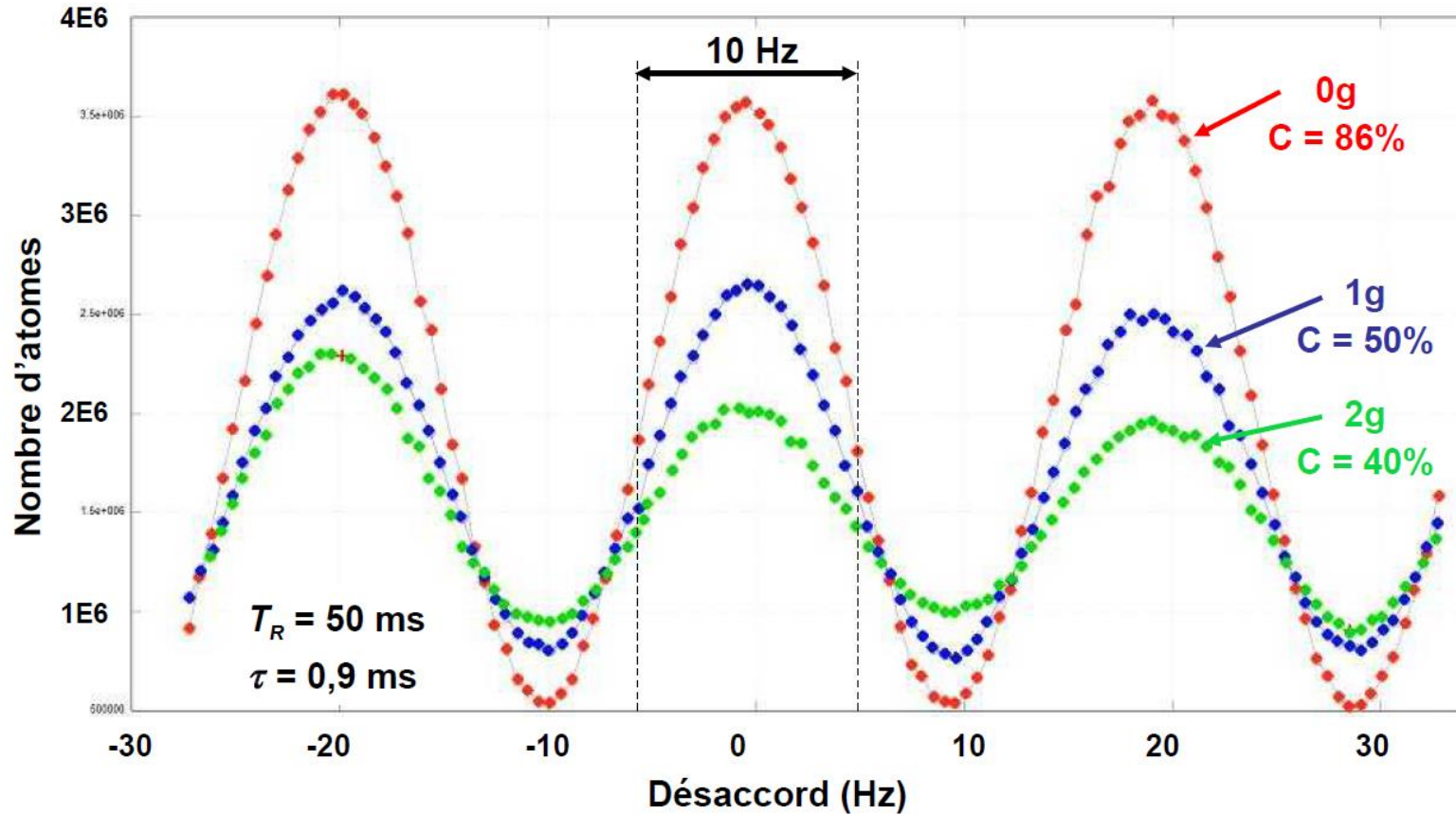
Possibilité de faire des mesures en « 0g », en « 1g » et en « 2g »



- Environnement accélérométrique relativement perturbé ($\sim 10^{-2}$ g suivant Z)
- Environnement magnétique très variable (jusqu'à 1G de variation entre l'entrée et la sortie de la parabole)

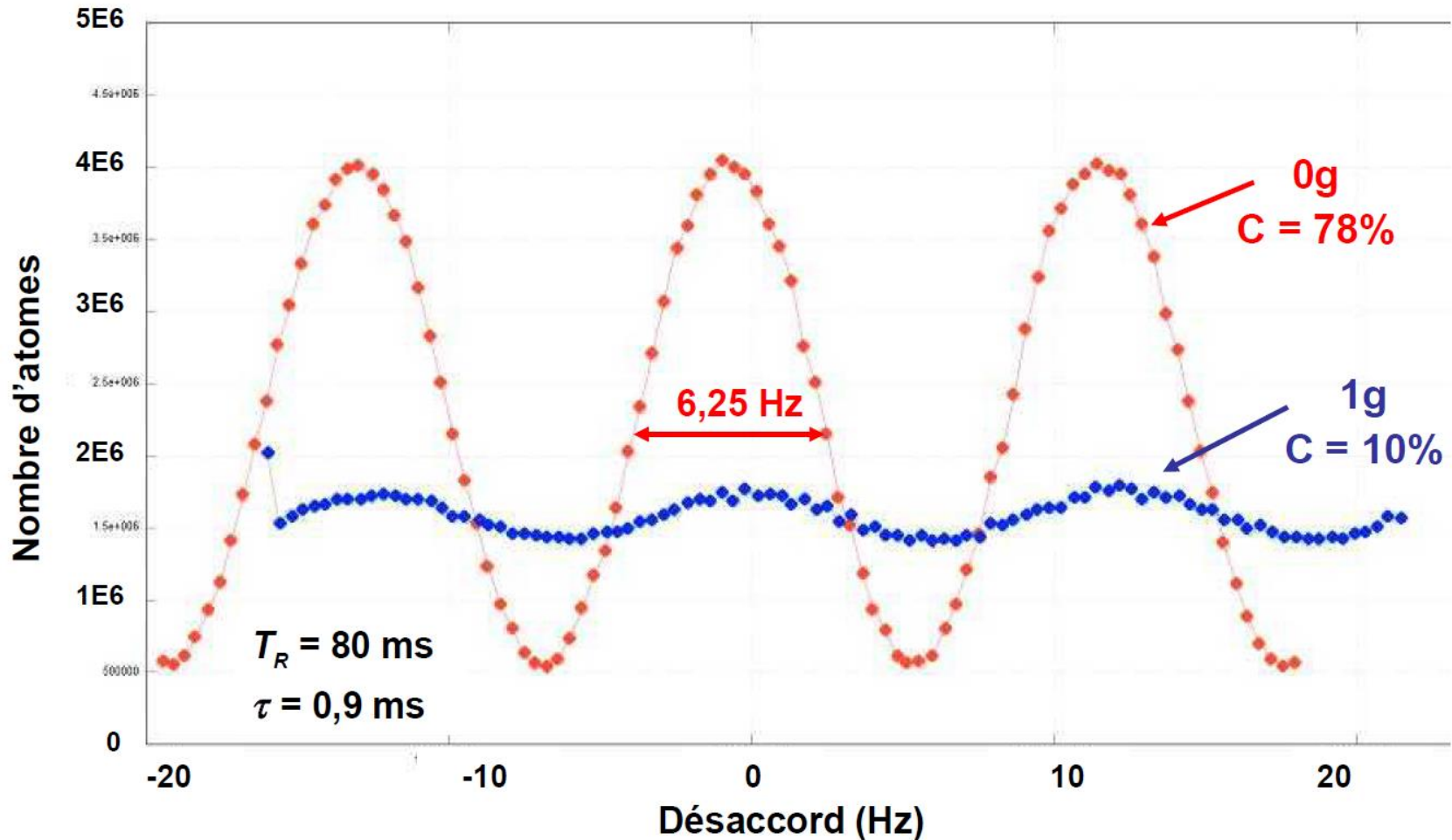
**3 vols de 31 paraboles chacun – 20 secondes de microgravité par parabole
30 minutes de mesure, par tranche de 20s, en tout et pour tout !!**

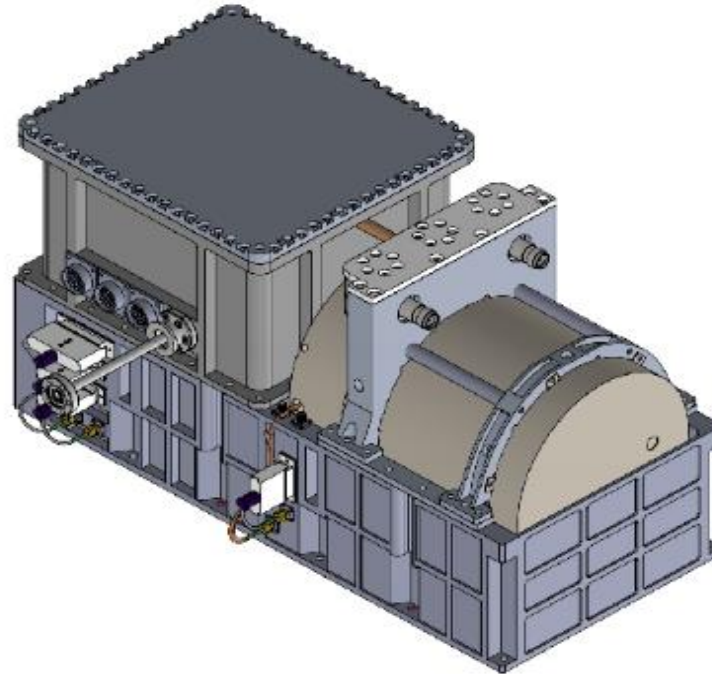
Comparaison entre « 0g », « 1g » et « 2g » pour un temps de ramsey de 50 ms



Gain d'un facteur 1,4 sur le nombre d'atomes et d'un facteur 1,7 sur le contraste pour $T_R = 50 \text{ ms}$

Comparaison entre « 0g » et « 1g » pour un temps de ramsey de 80 ms

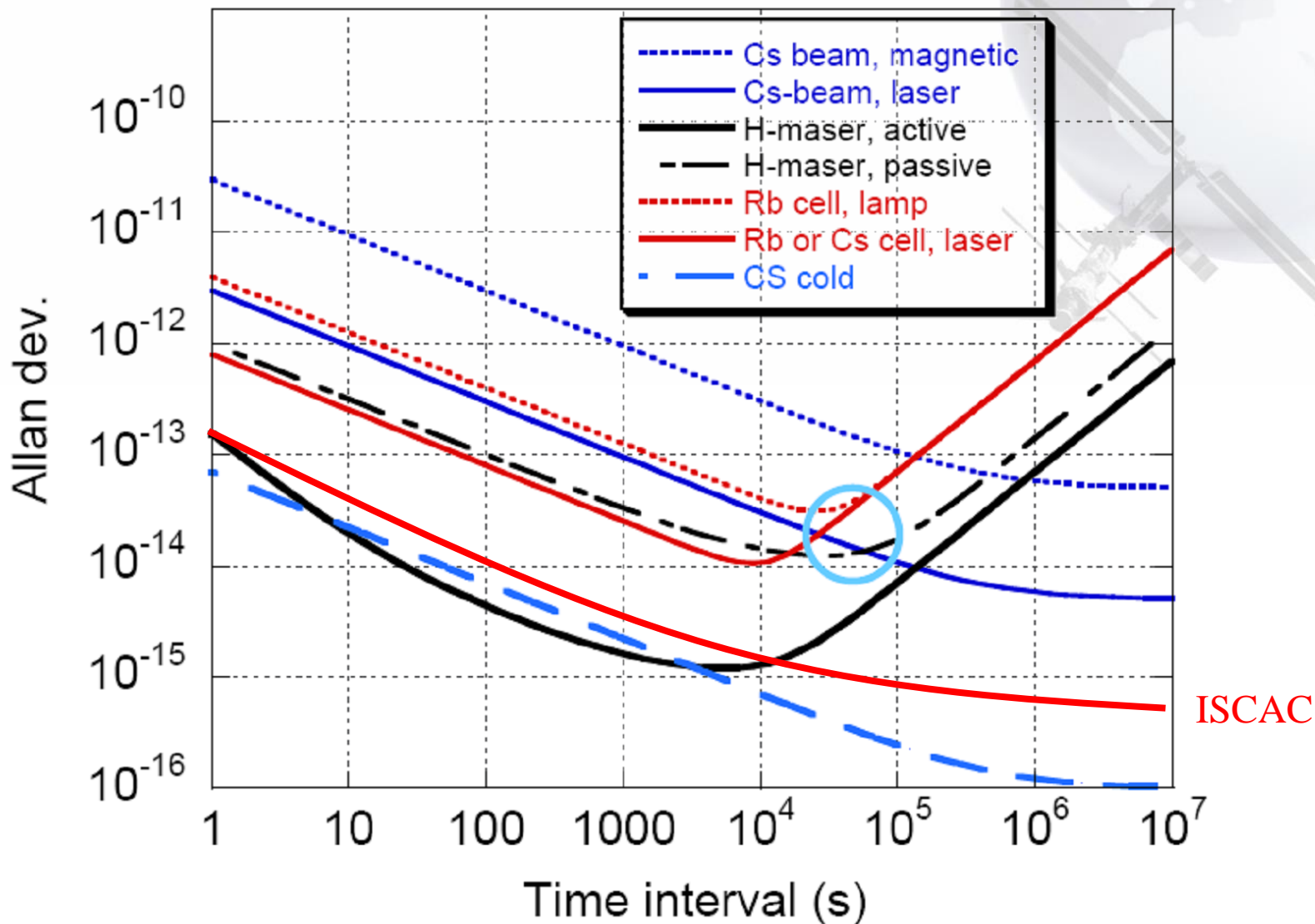




Flight model of ISCAC



Stability of atomic clocks



Thanks for attention

